BEEN SEEN BEFORE







SWG 19: Pasadena PLASTIC Status Report

Toni Galvin

M. Popecki, K. Simunac, B. Klecker L. Kistler, L. Ellis, C. Farrugia, E. Moebius, Y. Liu M. Lee, J. Barry, P. Bochsler, P. Wurz, L. Blush R. Wimmer-Schweingruber, A. Opitz, L. Berger

Solar Minimum:

Yesterday the Groundhog saw his shadow, confirming Chris StCyr's report:

Six more months of Solar Minimum!



Figure received from Alex Young

Personnel Changes

- Katherine Singer is now working at Hamilton Sunstrand (ISS Air Quality Control)
- Yong Liu, New Post Doc at UNH working on generation of Oxygen Parameters (Public Domain)
- Peter Bochsler, Visiting Scientist at UNH
- Submission of three candidates to NASA Project for Data Phase Co-Investigators: K. Simunac (UNH), A. Opitz (CESR), and L. Berger (CUA, Kiel)

Instrument Activities - 2008

Α

- 2/7/08 DPU FSW load 3.2.7
- 4/14/08 High voltages reset; instrument was restarted.
- 5/11/08 Instrument stalled and entered anomalous hv mode due to delivery timing of some memory register commands

Two previously unknown problems; both needed to create the hangup:

- Leap-year error in the POC s/w (since corrected by UCB, thanks!)
 +
- Variation in S/C from MOC-POC ICD about handling time-tag commands

Β

3/3/08 - DPU FSW load 3.2.7

11/19/08 MCP HV increase - 20V

11/21/08 Raised DPU-side S-ch switch threshold

11/26/08 MCP HV increase - 20V

A_RA_Trig Full Resolution



Science

In addition to these published papers cited in the bibliography, multiple papers have been submitted to Solar Physics, covering a variety of topics.

Others are in progress for the SOHO-STEREO Workshop.



STEREO Investigations



during the Recent Solar Minimum: **Observations and Consequences for Space** Weather Modeling

Applications of Grad-Shafranov Reconstruction Techniques on Magnetic Clouds Observed by Stereo and Wind and Comparison to Solar Source

Interstellar Pickup Ion Observations by STEREO: Focus Cone Variations with Solar Wind Streams

High Resolution Minor Ion Composition and Kinetic Properties at Solar Wind Interfaces --Solar Source Back Mapping

Suprathermal Ions in Compression Regions and in Upstream Events

STEREO Investigations (Solar Physics)

Is the Current Solar Cycle Minimum Weaker than Previous Cycles?

STEREO - Wind - Heiios Observations of the Solar Wind QI

The solar wind quasi-invarient (QI) is the ratio of the solar magnetic wind magnetic energy density to the plasma kinetic energy density, that is, the inverse square of the Alfven Mach number.

QI complements other indices and has implications on the Sun-Solar System connection during an entire solar activity phase.

On average, the QI(2007) is less than that observed during the previous two minima, indicating weaker solar activity. This also implies weaker MHD effects in solar wind flow around planetary magnetospheres which, in turn, alters the flow's interaction with them. Distribution of QI values in the year 2007 from STEREO-A, STEREO-B, and Wind. For comparison, we show the corresponding distributions at the minima of solar cycle 20 (Helios) and 22 (Wind).



Reference:

(1) M. Leitner, C. J. Farrugia, A. B. Galvin, H. K. Biernat, and V. A. Osherovich, The solar wind quasiinvariant observed by STEREO A and B at solar minimum 2007, and comparison with two other minima, *Solar Phys.*, under review, 2009.

STEREO Investigations (Solar Physics)

Temporal Evolution of the Solar Wind during the Recent Solar Minimum:

Observations and Consequences for Space Weather Modeling

The heliocentric orbits of the two STEREO satellites provide a unique opportunity to study the cross-correlation of solar wind parameters^{1,2}, and, during these solar minimum conditions, the evolution of stream interfaces³ near 1 AU over time scales of hours to a few days, that is, much less than the period of a Carrington rotation.

Presentation this SWG by Kristin Simuanc



References:

- (1) A. Opitz et al., Temporal evolution of the solar wind bulk velocity at solar minimum by cross-correlating the STEREO A and B PLASTIC measurements, *Solar Phys.*, accepted, 2009.
- (2) J.J. Podesta, A.B. Galvin, C.F. Farrugia, Correlation length of solar wind fluctuations measured tangent to the earth's orbit: First results from STEREO, *J. Geophys. Res.*, 113, A09 104, doi: 10.1029/2007JA012865, 2008.
- (3) K.D.C. Simunac, et al., In Situ Observations of Solar Wind Stream Interface Evolution, *Solar Phys.*, under review, 2009.

STEREO Investigations (Solar Physics)

Applications of Grad-Shafranov Reconstruction Techniques on Magnetic Clouds Observed by Stereo and Wind

Comparison to Solar Source

Multi-spacecraft STEREO and Wind observations allowed us to model two magnetic clouds (MCs) with increased accuracy. The model was a Grad-Shafranov Reconstruction where data from two spacecraft were ingested to optimize the magnetic field maps. The MCs were observed in May 2007 when the STEREO spacecraft where about 9° apart.

For one MC it was possible to compare the magnetic flux content with that inferred from the corresponding two-ribbon flare on the Sun. For the first time it was shown that most of the magnetic flux of the MC, now better constrained because of the available multispacecraft observations, is created during the eruption on the Sun. The shape of the magnetic clouds in a plane perpendicular to the invariant axis is composed of field lines showing a deformed elliptical shape (yellow sphere is the Sun, the spacecraft are from left to right: STEREO-B, WIND (Earth), STEREO-A).



References:

- Kilpua, E., P. Liewer, C. J. Farrugia, J. G. Luhmann, C. Moestl, et al., Multi-spacecraft observations of magnetic clouds and their solar origins: May 19123, 2007, *Solar Phys.*, in press, 2009.
- (2) C. Moestl, C. J. Farrugia, C. Miklenic, M. Temmer, A. B. Galvin, J. G. Luhmann, et al., Multi-spacecraft recovery of a magnetic cloud and its origin from magnetic reconnection on the Sun, *J. Geophys. Res.*, in press, 2009.
- (3) C. Moestl, C. J. Farrrugia, M. Leitner, E. K. J. Kilpua, A. B. Galvin, et al., Optimized reconstruction of a small magnetic cloud using STEREO-WIND observations, *Solar Phys.*, under review, 2009.

STEREO Investigations (in progress)

Interstellar Pickup Ion Observations by STEREO: Focus Cone Variations with Solar Wind Streams

Pickup helium ions provides diagnostics for Local Interstellar Cloud parameters, such as the LIC flow velocity vector and temperature. However, pickup ion distributions are highly variable and are known to be affected by the interplanetary magnetic field, solar wind density, ionization rates & probably several unknown causes. For deduction of LIC parameters, mitigation is sought with better information to distinguish temporal vs spatial differences. STEREO A/B provide two traversals of the focusing cone per year.

STEREO RESULTS:

Presentation this SWG by Berndt Klecker







References:

(1) B. Klecker et al., AGU 2008, SWG 2009

(2) E. Moebius, B. Klecker et al. Simultaneous
Pickup Ion Observations of the Interstellar Helium
Focusing Cone
with STEREO PLASTIC A
& B, Cospar 2008.



Data available through SSC web-site:

- All Level 1 CDF: Housekeeping, monitor rates, and all science data, including proton onboard moments. Does not include heavy ion efficiencies. (These files are used in analysis programs SPLAT and PHA_PLAY.)
- Level 2 CDF 1-minute cadence data for the 1D-Maxwellian derived proton bulk parameters. March 2007 through Nov. 2008. These files lag by about 30 days because they require the *.fin Level 0 files.

Data: epoch, epoch_1kev, error codes, caution codes, s/c attitude flag, proton density, proton bulk speed, kinetic temperature, thermal velocity, n_s_flow_angle, carrington_rotation, spcrft_lon_carr, heliospheric_dist, spcrft_lon_hee, spcrft_lat_hee, spcrft_lon_heeq, spcrft_lat_heeq, spcrft_lon_hci, spcrft_lat_hci. New STA processing this week, to incorporate E_W flow angle, and velocity components (RTN, HERTN).







Data available through PLASTIC web-site

Ascii/Excel Proton bulk parameters: 1 min (4/2007-9/2008), 10 min (2/2007-9/2008), 1 hr (2007-7/2008)

Automatically Generated Plots:

Level 2 Proton Moments (1D Max) Bulk Vel, Density, Temperature, and Thermal_vel through present

On-Board Moments: Bulk Vel, Density, Temperature, Vel. Standard Deviation through present

Solar Wind Stack Plots through 10/2008 He+ Survey Plots through present Spacecraft roll, pitch, and yaw through

present

Movie: Solar Wind Fe Charge States at CIR Coming this week: STA Proton Velocity Components

Upcoming Meetings

- STEREO SWG for Fall
- SOHO CELIAS / PLASTIC STEREO Team meeting April 7-9 2009, to be held in Concord NH at the STEREO EPO Partner McAuliffe-Sheppard Discovery Center
 - •If interested in attending, please contact
 - •<u>toni.galvin@unh.edu</u> and berndt.klecker@mpe.mpg .de

